

Quantum Fluctuations in Simplicial Gravity

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- 2 Causal Dynamical Triangulations
- 3 Background spacetime
- 4 The Minisuperspace Model
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Path integral formulation of Quantum Gravity

- The partition function of quantum gravity is defined as a formal integral over all geometries weighted by the Einstein-Hilbert action.
- To make sense of the gravitational path integral one uses the standard method of regularization - discretization.
- The path integral is written as a nonperturbative sum over all causal triangulations \mathcal{T} .
- Wick rotation is well defined due to proper time foliation.
- Using Monte Carlo techniques we can calculate expectation values of observables.

$$Z = \int \frac{\mathcal{D}_{\mathcal{M}}[g]}{\text{Diff}_{\mathcal{M}}} e^{iS^{EH}[g]} \quad \rightarrow \quad Z = \sum_{\mathcal{T}} \frac{1}{s(\mathcal{T})} e^{iS^{\text{Regge}}[g]}$$

Causal Dynamical Triangulations (CDT) is a background independent approach to quantum gravity.

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Path integral formulation of Quantum Gravity

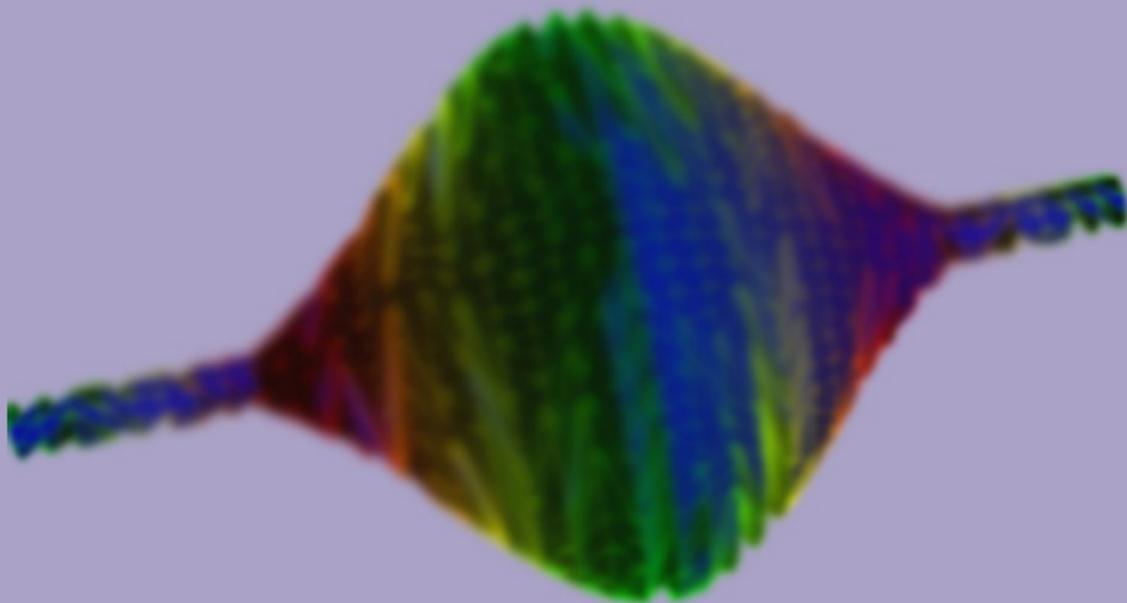
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Causal Dynamical Triangulations (CDT) is a background independent approach to quantum gravity.

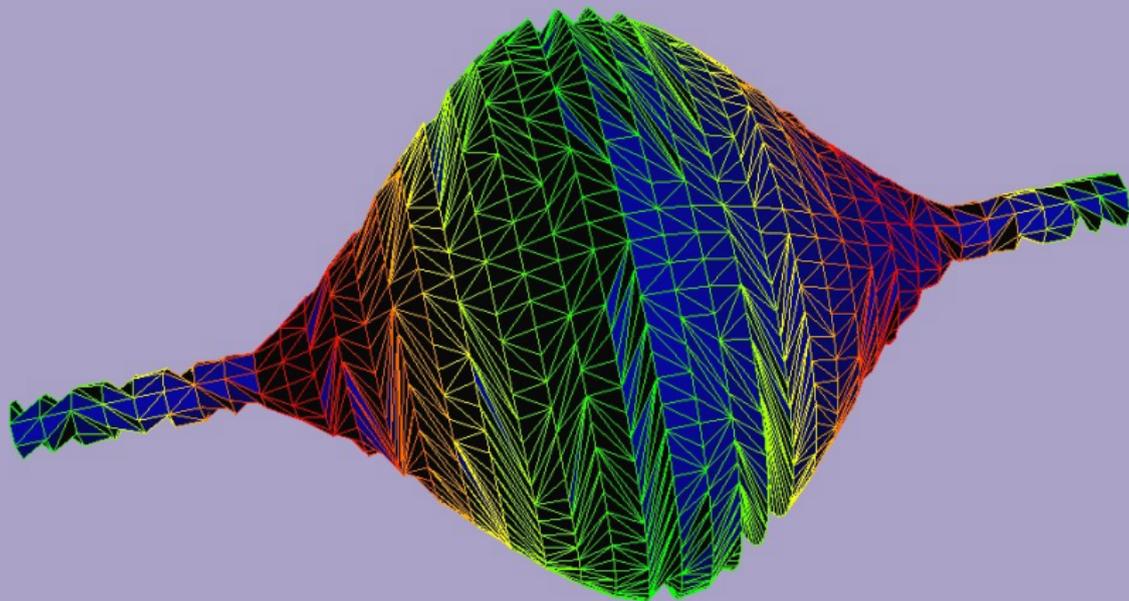
Dynamical Triangulations

A manifold with topology $S^3 \times S^1 \dots$



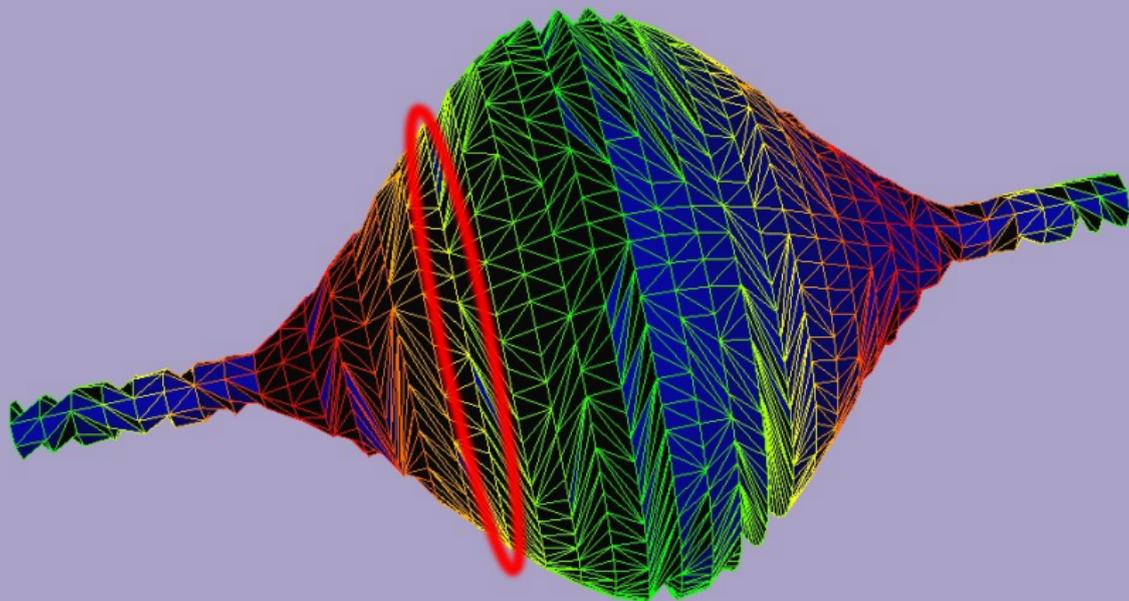
Dynamical Triangulations

... is discretized by gluing 4-simplices - triangulation



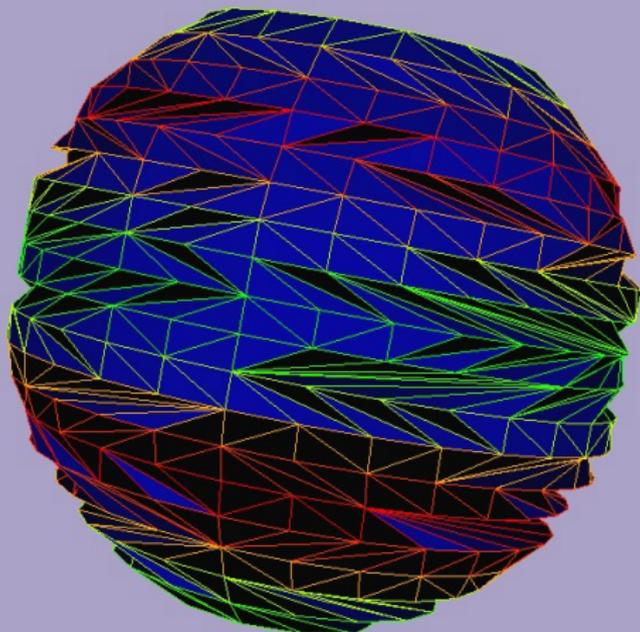
Causal Dynamical Triangulations

The spatial slices (of constant time) . . .



Causal Dynamical Triangulations

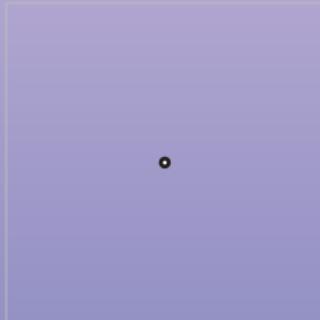
... have S^3 topology



Fundamental building blocks of CDT

- d -dimensional simplicial manifold can be obtained by gluing pairs of d -simplices along their $(d - 1)$ -faces.
- Lengths of the time and space links are constant. Simplices have a fixed geometry.
- The metric is flat inside each d -simplex.
- The angle deficit (curvature) is localized at $(d - 2)$ -dimensional sub-simplices.

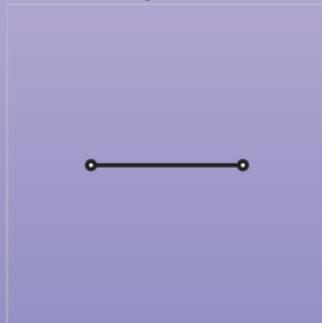
0D simplex - point



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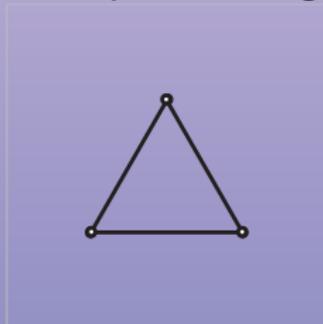
1D simplex - link



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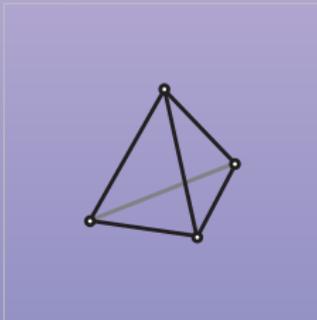
2D simplex - triangle



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3D simplex - tetrahedron



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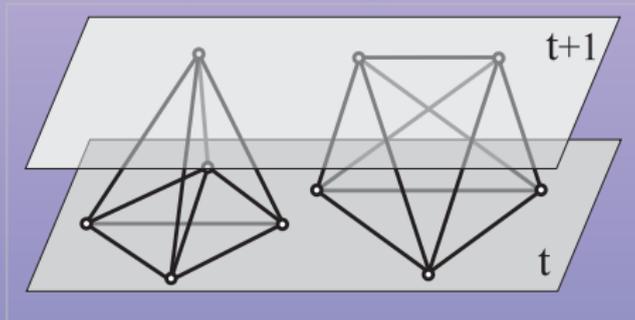
4D simplex - 4-simplex



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4D simplex, two types in CDT



Monte Carlo simulations - Alexander moves

We construct a starting space-time manifold with given topology ($S^3 \times S^1$) and perform a random walk over configuration space.

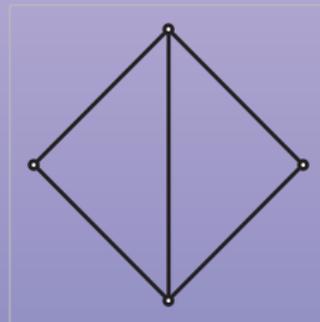
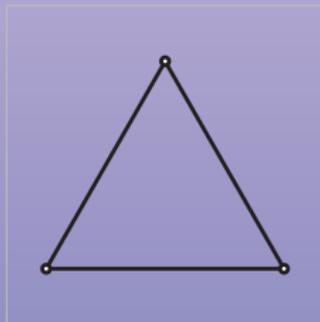
Ergodicity In the dynamical triangulation approach all possible configurations are generated by the set of Alexander moves.

Fixed topology The moves don't change the topology.

Causality Only moves that preserve the foliation are allowed.

4D CDT We have 4 types of moves.

Examples in 2D



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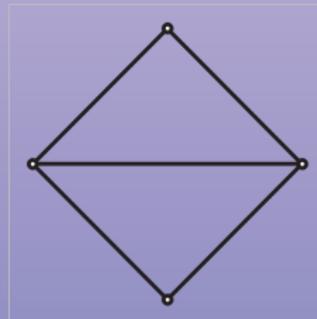
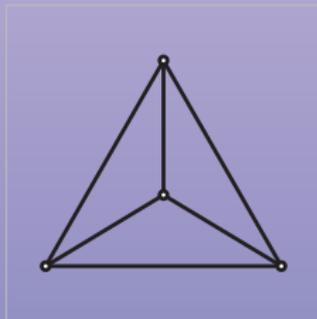
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Causal Dynamical Triangulation - properties

- Manifolds are approximated by simplicial manifolds.
- Sum over triangulations (gluings).
- By construction we have foliation in proper time. We do not allow the spatial slices to change topology.
- Wick rotation is well defined.
- Such formulation involves only geometric invariants like length and angles.
- We don't introduce coordinates.
- Manifestly diffeomorphism-invariant.

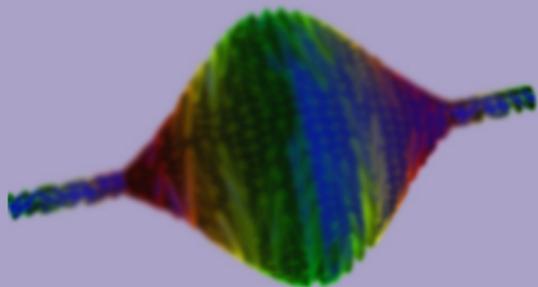
Path integral - The action

We generate a large number of such configurations with the probability

$$P[\text{configuration}] \propto e^{-S}$$

We use Einstein-Hilbert action ...

$$S = -\frac{1}{G} \int dt \int d\Omega \sqrt{g} (R - 6\lambda)$$



G gravitational constant

λ cosmological constant

g determinant of a spacetime metric

R scalar curvature

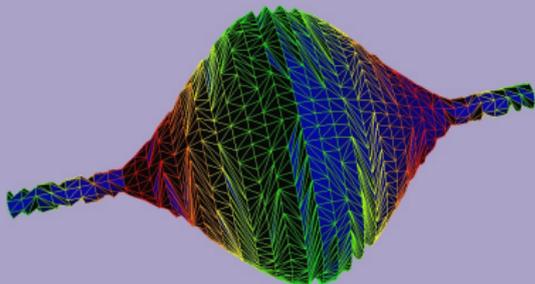
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...or the Regge action in the discrete version

$$S = -K_0 N_0 + K_4 N_4 + \Delta(N_{14} - 6N_0)$$



N_0 number of vertices

N_4 number of simplices

N_{14} number of simplices of type
 $\{1, 4\}$

K_0 K_4 Δ bare coupling constants

Background

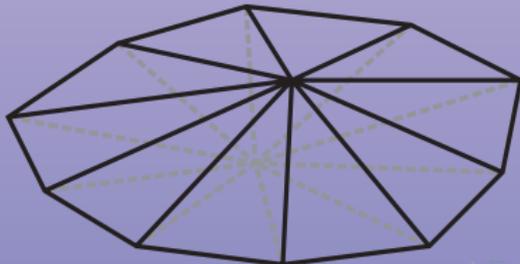
- For a certain range of bare coupling constants, a typical configuration has a "bloblike" shape and behaves as a well defined four-dimensional manifold.
- This isn't trivial. In the Euclidean version, without imposed causality, one either got
 - a "crumpled phase" with infinite Hausdorff dimension
or
 - a "branched polymer phase" dominated by spacetimes where the 4-simplices form treelike structures with Hausdorff dimension two,

even though they are built from 4D simplices. Unfortunately, the phase transition between them is of the 1st order.

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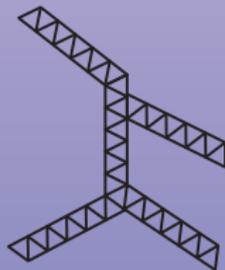
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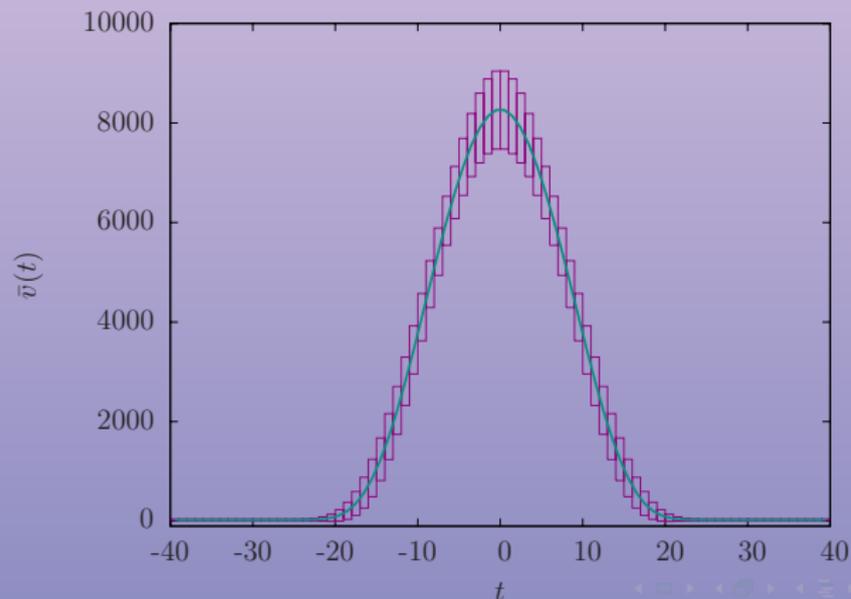
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Background

- For a certain range of bare coupling constants, a typical configuration has a "bloblike" shape and behaves as a well defined four-dimensional manifold.
- The averaged spatial volume $\bar{v}(t)$ is proportional to $\cos^3(t/B)$.



Background - classical solution

- Such behaviour of the average volume occurs when we assume spatial homogeneity and isotropy.
- We "freeze" all degrees of freedom except the volume (scale factor).
- We introduce the following metric on $S^3 \times S^1$ spacetime

$$ds^2 = dt^2 + v^{2/3}(t)d\Omega_3^2$$

- In this particular case, the Einstein-Hilbert action

$$S = \frac{1}{G} \int dt \int d\Omega \sqrt{g}(R - 6\lambda)$$

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- In this particular case, the Einstein-Hilbert action takes form

$$S = \frac{1}{G} \int \frac{\dot{v}^2}{v} + v^{\frac{1}{3}} - \lambda v dt$$

The classical trajectory $\bar{v}(t) \propto \cos^3(t/B)$ is perfectly recovered in this model.

Question?

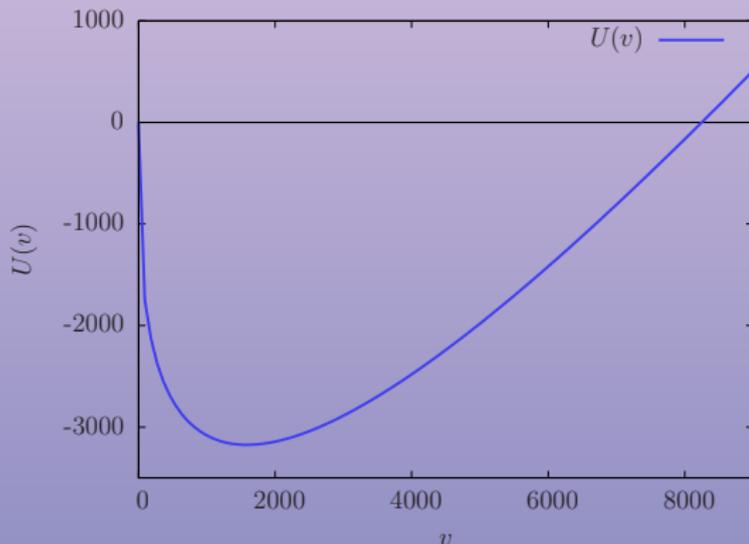
How well does the **minisuperspace** model describe the quantum fluctuations computed from Monte Carlo simulations?

The **minisuperspace** action

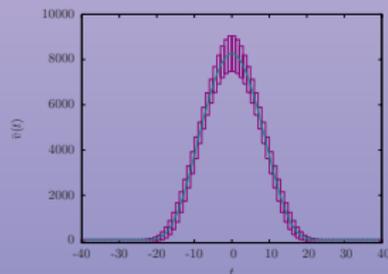
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The potential

- Restricting our considerations to the volume $v(t)$ we reduce the problem to one-dimensional quantum mechanics.
- The **minisuperspace** action describes a motion of a particle in a potential:

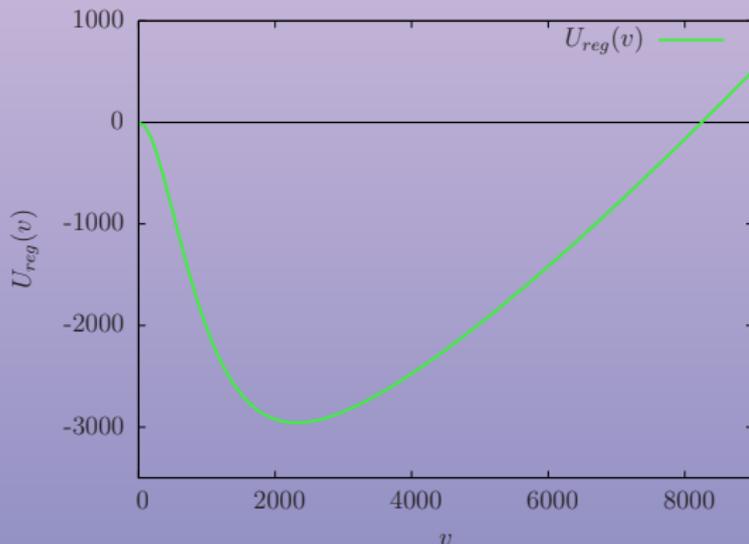


$$U(v) = -v^{\frac{1}{3}} + \lambda v$$

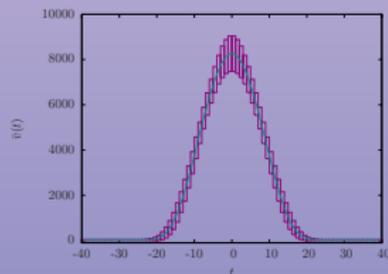


The regularized potential - bounce

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Quantum fluctuations - semiclassical approximation

- The spatial volume fluctuations are described by a Hermitian Sturm-Liouville operator $D(t)$.

$$S[v = \bar{v} + x] \approx S[\bar{v}] + \frac{1}{2} \int x(t) D(t) x(t) dt$$

$$D(t) = -\partial_t \frac{1}{\bar{v}(t)} \partial_t + \left. \frac{\partial^2 U}{\partial v^2} \right|_{v=\bar{v}}$$

- For a discrete time it is a matrix $M_{tt'}$

$$S[v = \bar{v} + x] \approx S[\bar{v}] + \frac{1}{2} \sum_{t,t'} x_t M_{tt'} x_{t'}$$

$$x_t M_{tt'} x_{t'} = c_1 \sum_t \frac{(x_{t+1} - x_t)^2}{\bar{v}_t} + \left. \frac{\partial^2 U}{\partial v^2} \right|_{v=\bar{v}} x_t^2$$

- Can we recover the matrix $M_{tt'}$ from simulations?

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Propagator vs action

- There exists a direct relationship between the propagator - fluctuation correlation matrix - and the matrix M

$$C_{tt'} = \langle x_t x_{t'} \rangle = \frac{1}{Z} \int x_t x_{t'} e^{-\frac{1}{2} \sum_{t,t'} x_t M_{tt'} x_{t'}} \prod_s dx_s = M^{-1}_{tt'}$$

- The propagator C can be measured. $x_t = v_t - \bar{v}_t$.
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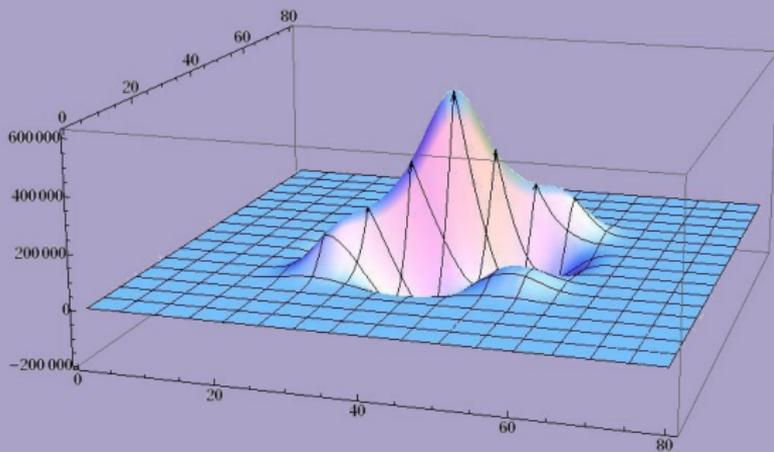
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The propagator C

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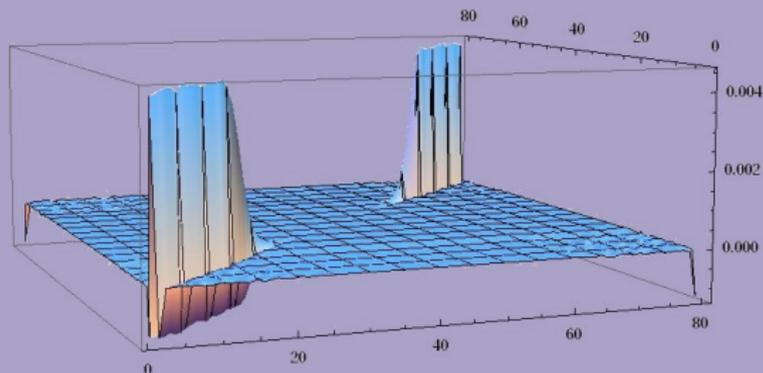
How does it look like?



The Sturm-Liouville operator M

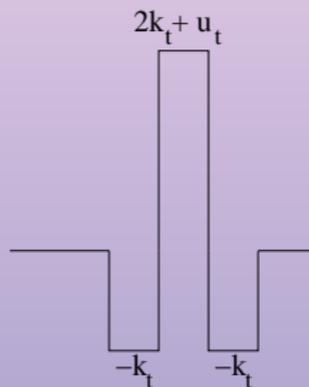
$$M = C^{-1}$$

How does it look like?



M matrix decomposition

- Analysing the matrix M row by row we can find the kinetic k_t and potential u_t coefficients.
- This allows us to compare the theoretical predictions with numerical data.
- We expect that $\bar{v}_t \approx \frac{c_1}{k_t}$ and $u_t \approx -U''(\bar{v}_t)$.

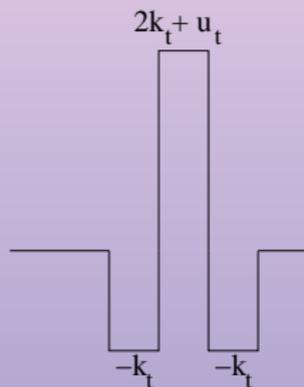


$$x_t M_{tt'} x_{t'}$$

$$c_1 \sum_t \frac{(x_{t+1} - x_t)^2}{\bar{v}_t} - U''(\bar{v}_t) x_t^2 = \sum_t k_t (x_{t+1}^2 + x_t^2 - 2x_t x_{t+1}) + u_t x_t^2$$

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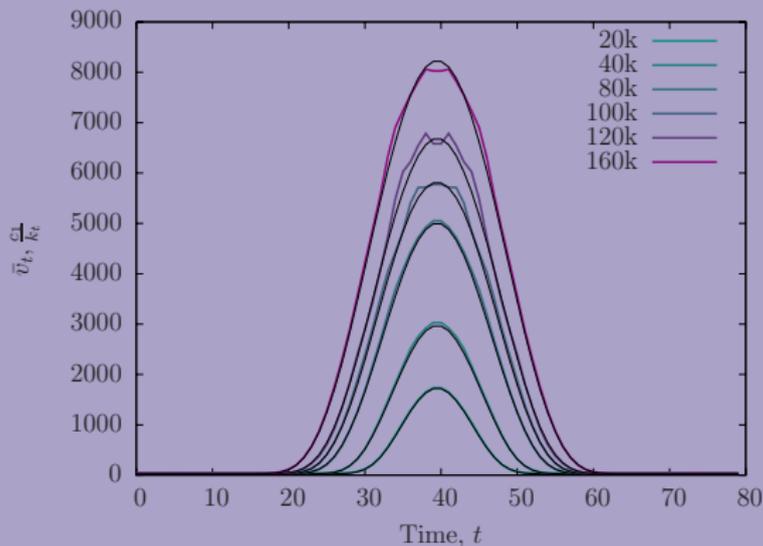
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The kinetic part

- Coefficients k_t fully agree with the predictions

$$\bar{v}_t = \frac{c_1}{k_t}$$

The kinetic part



The kinetic part

- Coefficients k_t fully agree with the predictions

$$\bar{v}_t = \frac{c_1}{k_t}$$

- The constant c_1 doesn't depend on the total volume.
- We can relate the **cut-off** with the gravitational constant G , which is responsible for the fluctuation amplitude.

$$G = \text{const} \frac{a^2}{c_1}$$

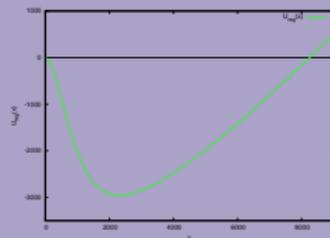
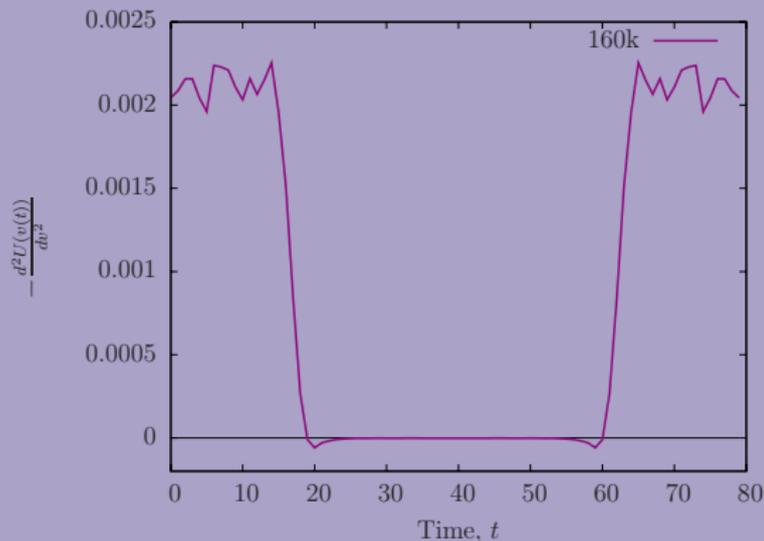
- The Universe built of 362000 simplices has a radius of about 20 Planck lengths.

The potential part

- The coefficients u_t also agree with the regularized potential

$$u_t \approx -U''(\bar{v}_t)$$

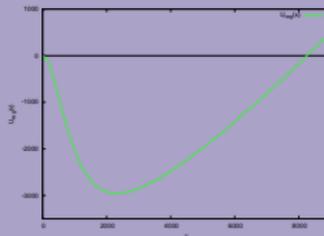
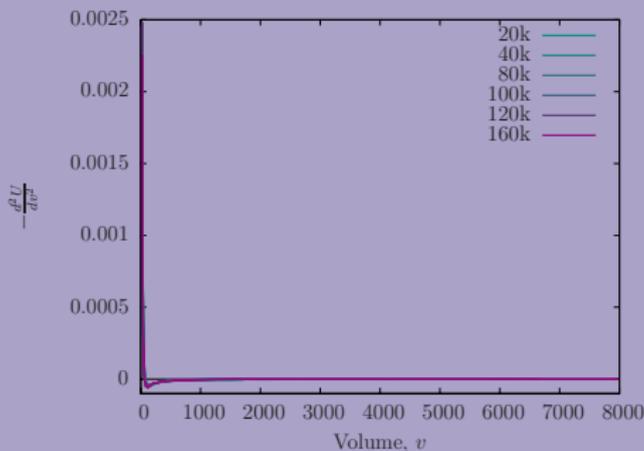
Potential well - $U''(v(t))$ plot



The potential part

- If the model is correct, $U''(v)$ should be universal and not depend on total volume. The "cosmological constant" λ controls the volume, but it gives no contribution to $U''(v)$.
- The numerical results are in full agreement with the regularised potential.

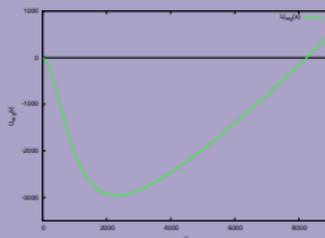
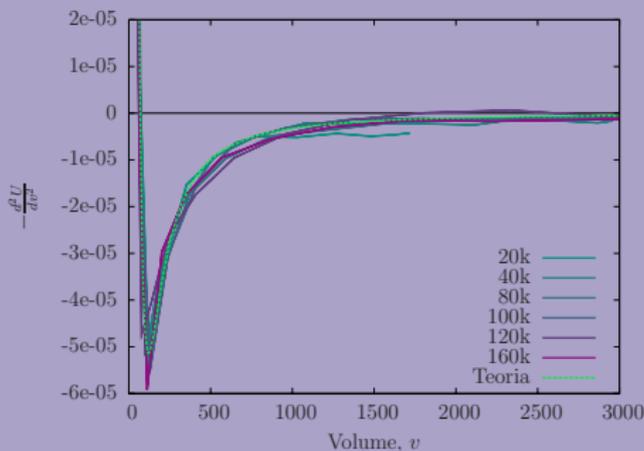
$U''(v)$ plot



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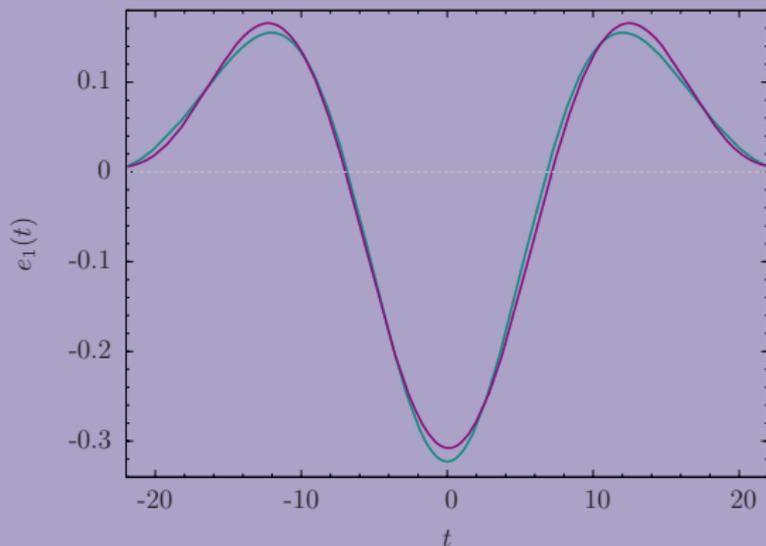
$U''(v)$ plot



Eigenvectors

- Eigenvectors of the numerical matrix C and theoretical matrix M are very similar.

The eigenvectors of C i M



Conclusions

- 1 We observe a four-dimensional universe with well defined time and space extension.
- 2 The background geometry exactly corresponds to the classical solution of the minisuperspace model (classical Einstein theory).
- 3 Quantum fluctuations are properly described by this simple model.
- 4 The gravitational constant G controls the fluctuation amplitude. We may estimate that the Universe built of 362000 simplices has a radius of about 20 Planck lengths.
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